

1 *IEEE Transactions on Geoscience and Remote Sensing*

2 **Supplementary Materials for**

3 **Towards Long-Wavelength Ionospheric Correction of InSAR Time Series**

4 **Using GNSS-Based TEC**

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17 **Content of this file**

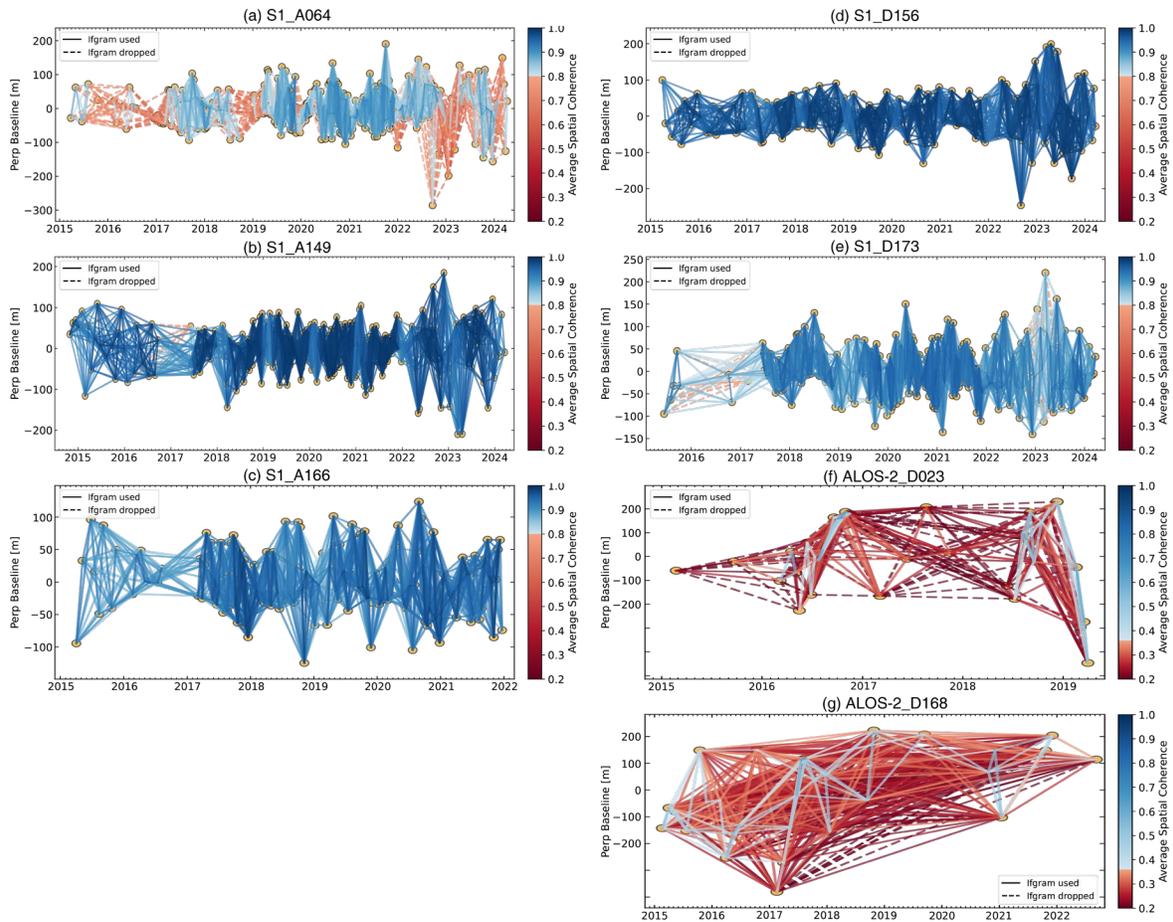
18 Section S1. Supplemental figures

19 Section S2. Sentinel-1 and ALOS-2 InSAR uncertainty due to orbit error

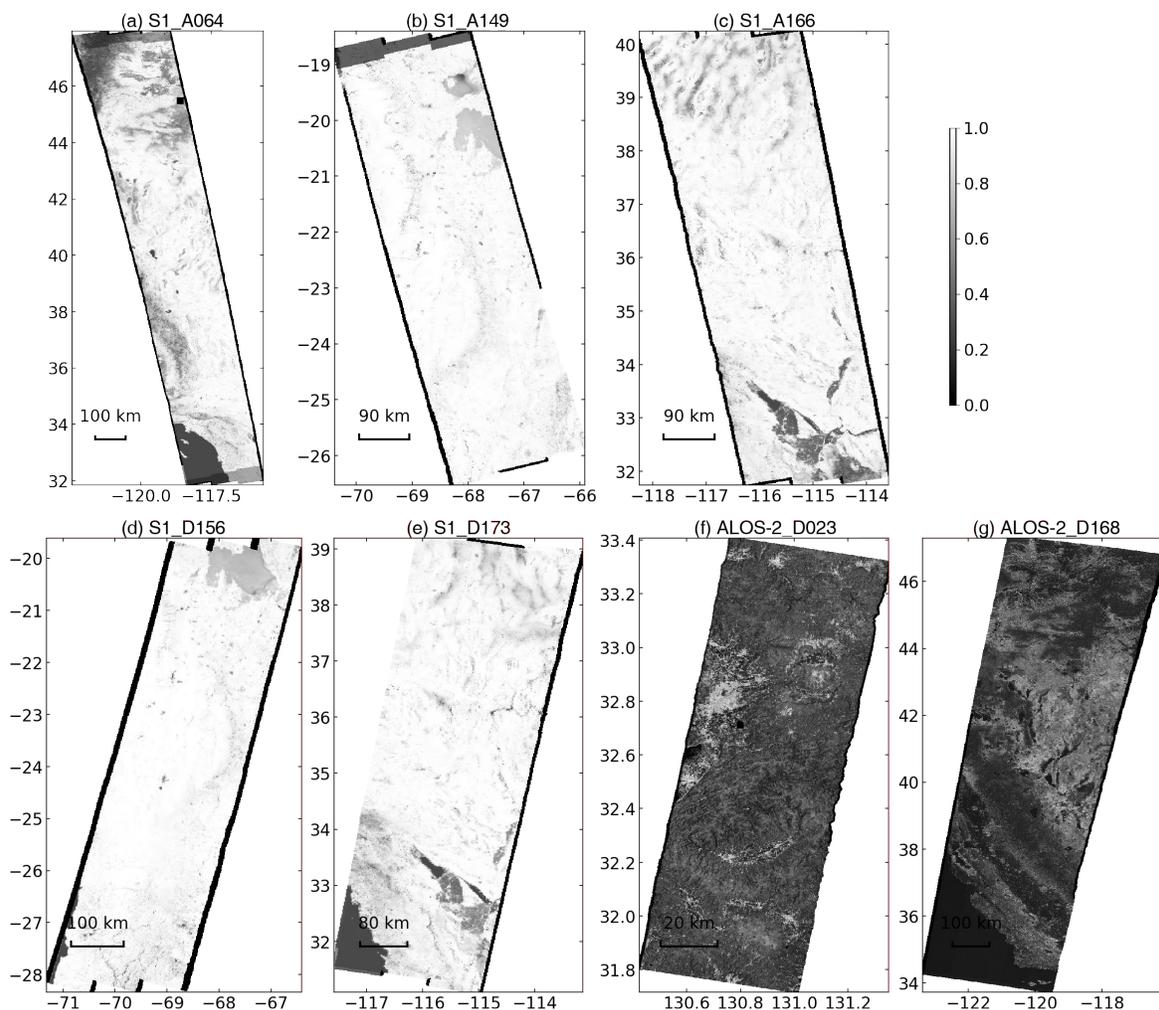
20 Section S3. Displacement accuracy evaluation using independent GNSS displacements

## 21 S1. Supplemental figures

22 This section provides Figures S1 to S10. Fig. S1 shows the interferogram networks of all  
 23 interferograms used in this study. Fig. S2 shows the average value of the spatial coherence of  
 24 all interferograms used in this study. Fig. S3 shows the temporal coherence of all  
 25 interferograms used in this study. Figs. S4-S10 show the different phase time series for Sentinel-  
 26 1 ascending track 166 in western USA.



27  
 28 **Figure S1.** The interferogram network of all Sentinel-1 and ALOS-2 tracks. Solid lines represent the  
 29 interferograms used for subsequent network inversion to estimate the time series, while dashed lines  
 30 represent the interferograms discarded due to low spatial coherence.



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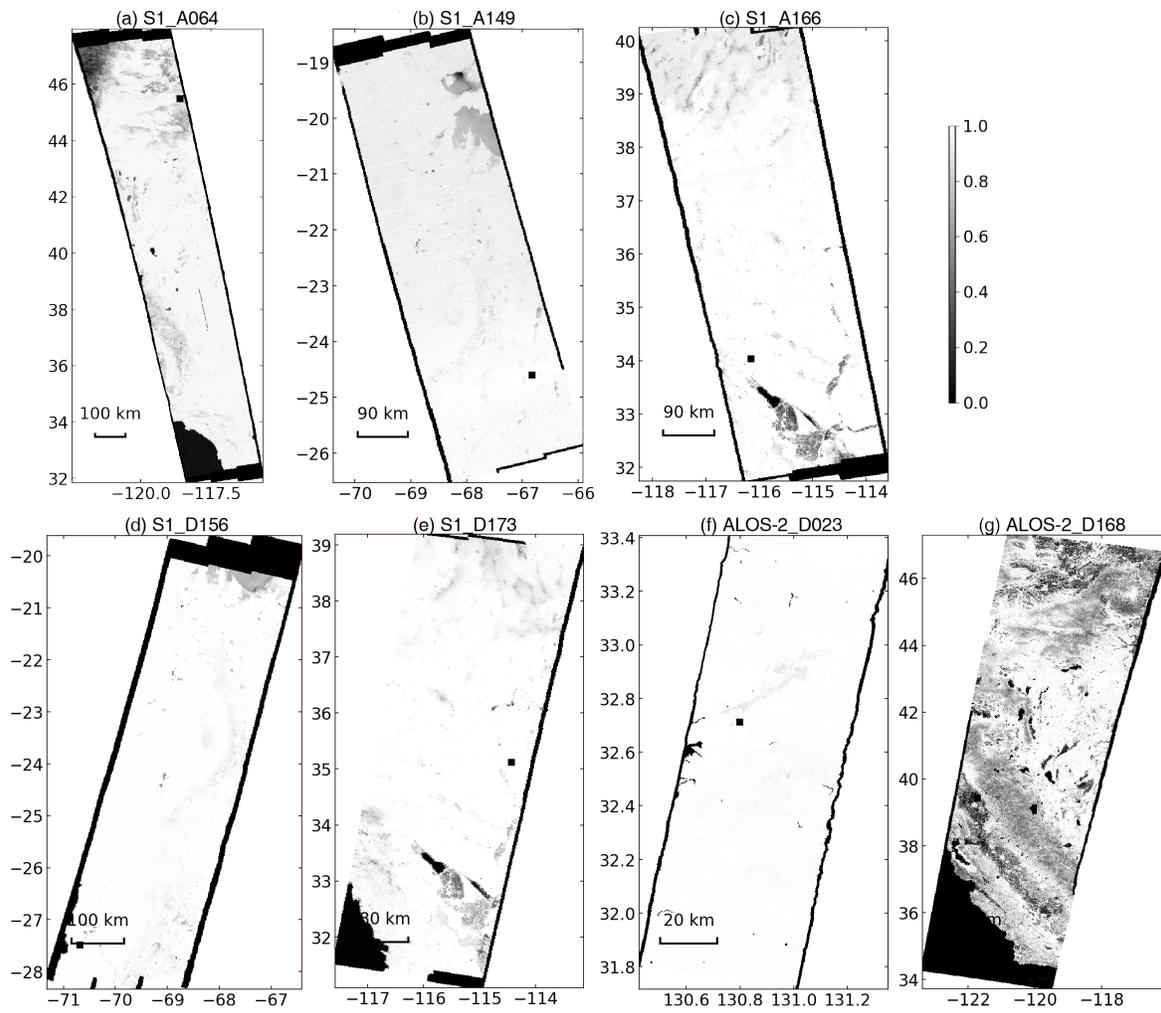
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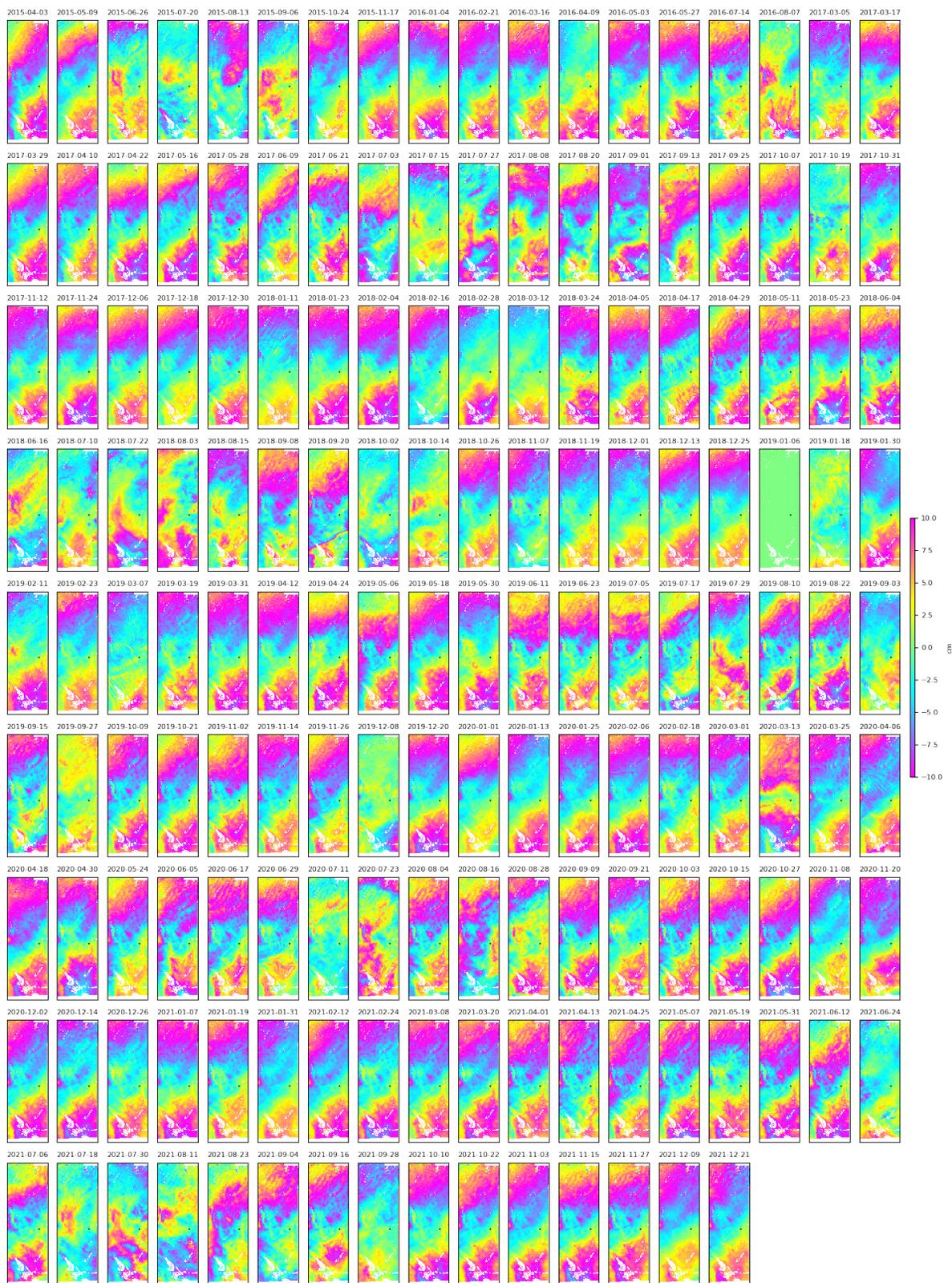
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**Figure S2.** The average spatial coherence of all Sentinel-1 and ALOS-2 tracks. The spatial coherence of the C-band Sentinel-1 (a)-(e) is calculated from the phase standard deviation using equation (12) in Agram & Simons (2015) [1], while the spatial coherence of the L-band ALOS-2 (f)-(g) is calculated using the complex coherence from SLCs using equation (11) in Agram & Simons (2015) [1].



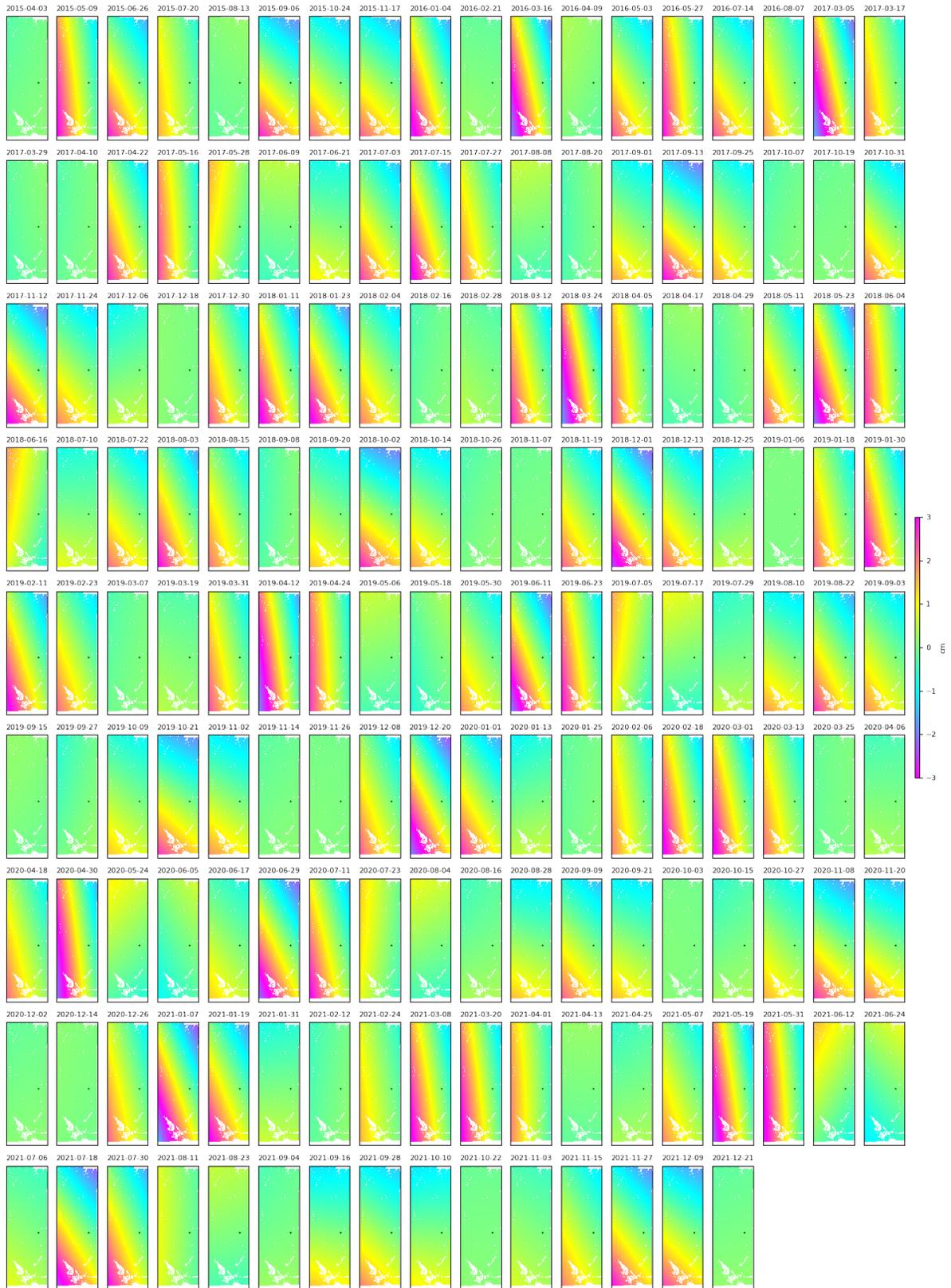
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37 **Figure S3.** The temporal coherence of all Sentinel-1 and ALOS-2 tracks. The temporal coherence is  
 38 calculated using equation (16) in Pepe et al. (2006) [2].



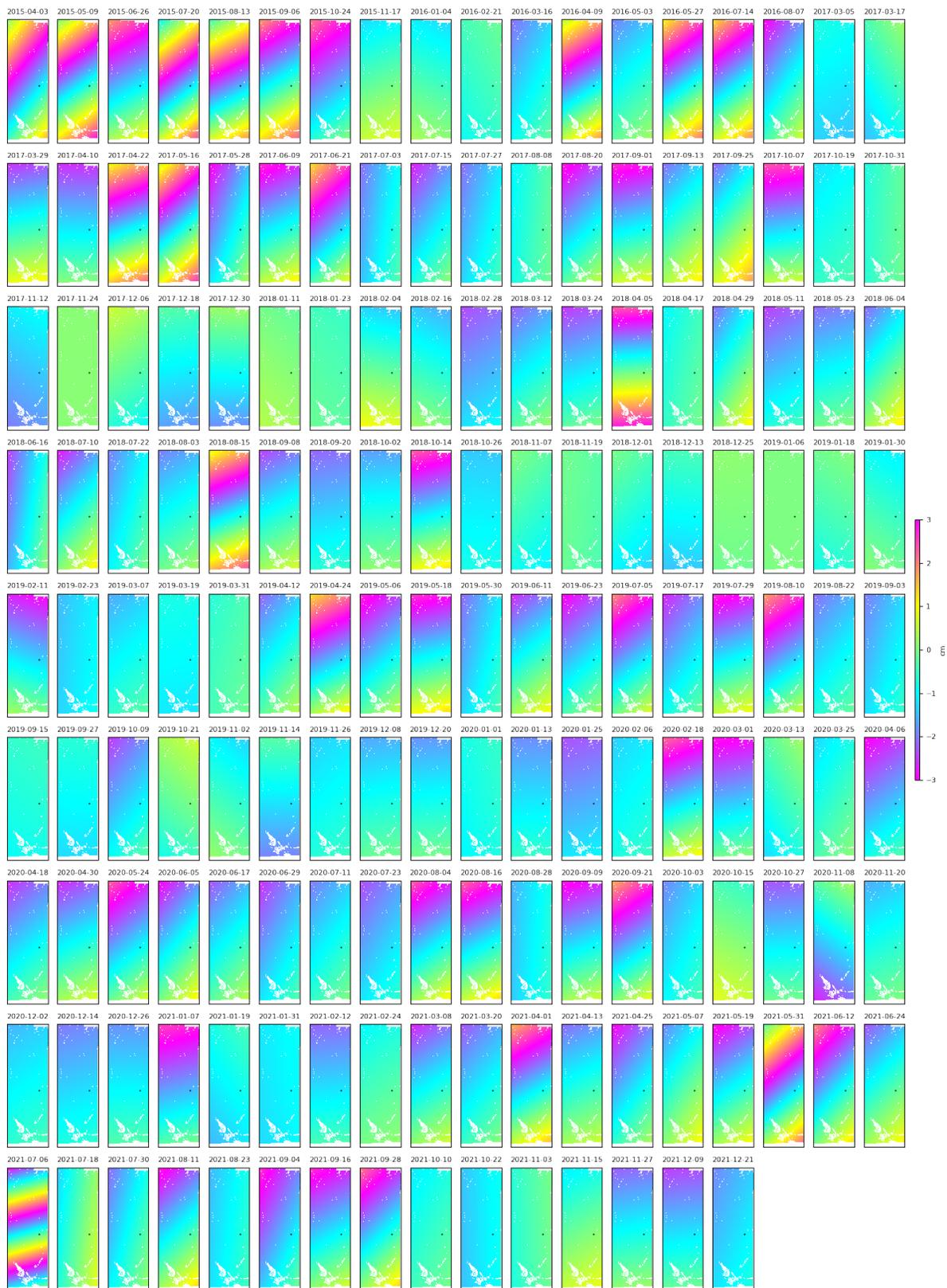
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**Figure S4.** The raw phase time series of Sentinel-1 ascending track 166 in western USA. The data is rewrapped into  $[-10, 10)$  cm for display. Black square donates the spatial reference point at  $35.8^{\circ}\text{N}$ ,  $114.9^{\circ}\text{W}$ .



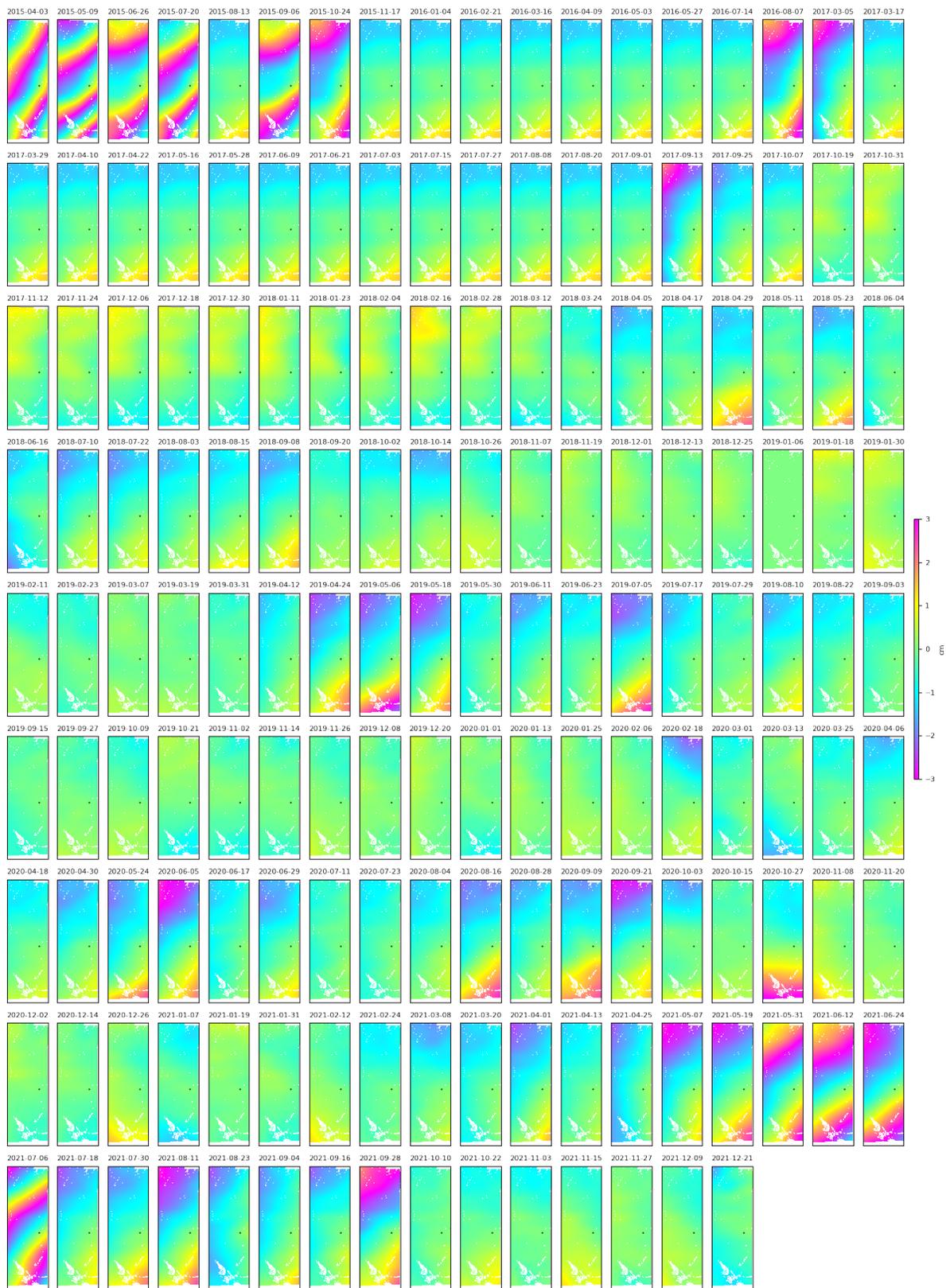
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**Figure S5.** Similar to Fig. S4, but for the solid Earth tides phase time series. The data is rewrapped into  $[-3, 3)$  cm for display.

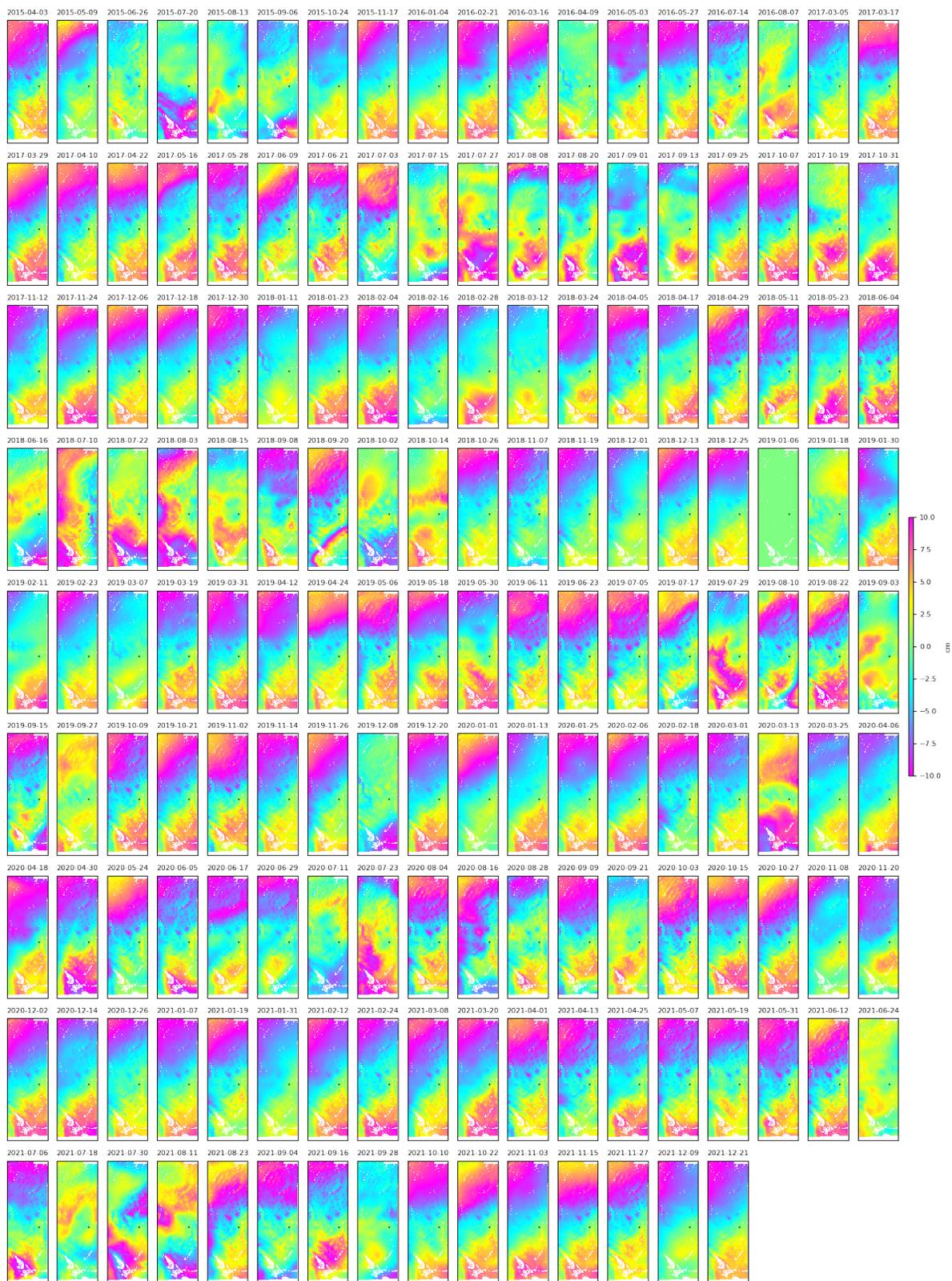


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**Figure S6.** Similar to Fig. S4, but for the ionospheric delay phase time series using Madrigal TEC products. The data is rewrapped into  $[-3, 3)$  cm for display.



**Figure S7.** Similar to Fig. S4, but for the ionospheric delay phase time series using range split-spectrum method. The data is rewrapped into  $[-3, 3)$  cm for display.



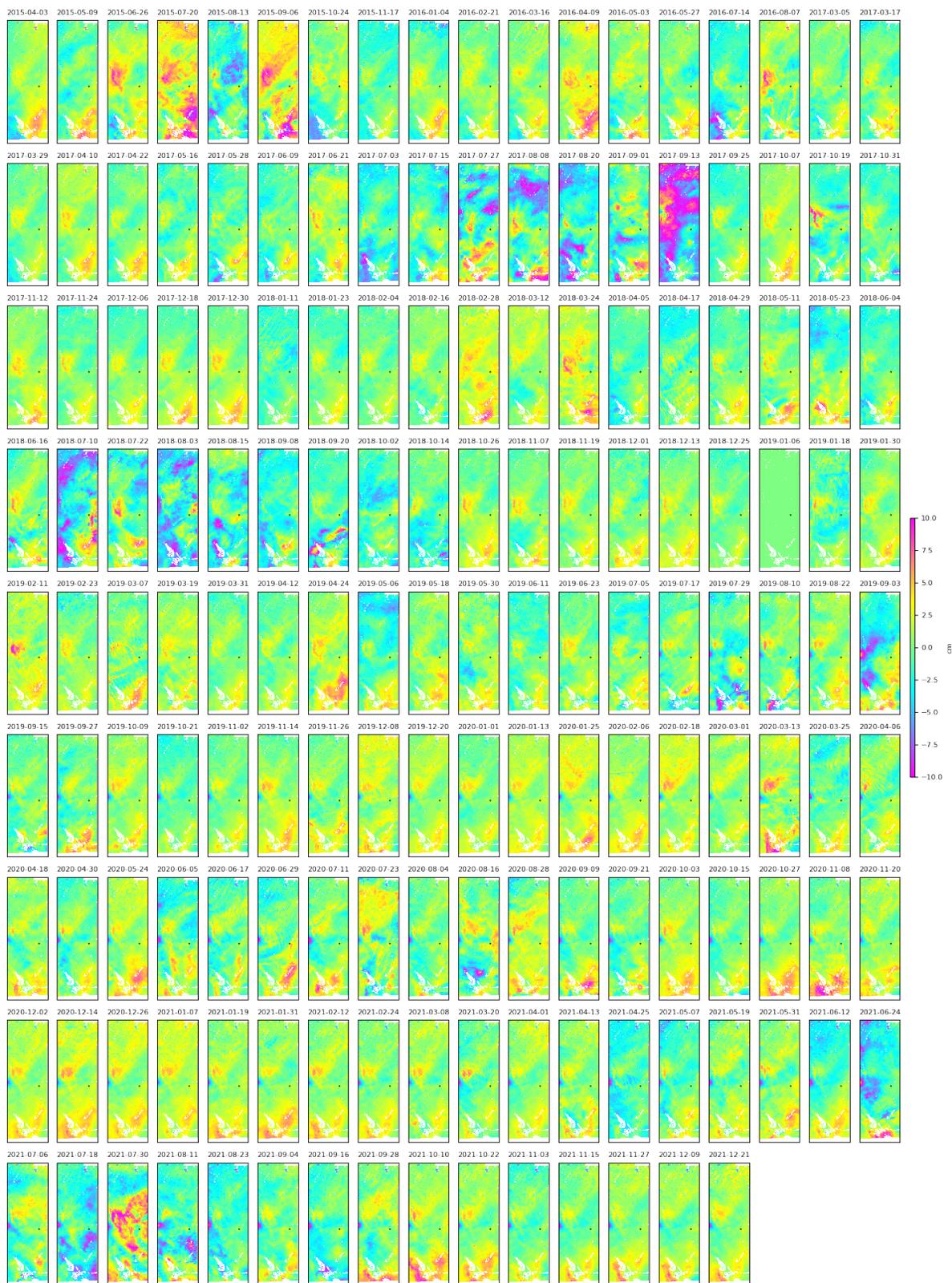
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**Figure S8.** Similar to Fig. S4, but for the tropospheric delay phase time series using the ERA5 weather reanalysis data. The data is rewrapped into  $[-10, 10]$  cm for display.



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**Figure S9.** Similar to Fig. S4, but for the topographic residuals phase time series from DEM errors. The data is rewrapped into  $[-1, 1)$  cm for display.



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**Figure S10.** Similar to Fig. S4, but for the final displacement time series after correcting for solid Earth tides (Fig. S5), ionosphere (using Madrigal TEC products; Fig. S6), troposphere (Fig. S8), and topographic residuals (Fig. S9). The data is rewrapped into  $[-10, 10]$  cm for display.

## 46 **S2. Sentinel-1 and ALOS-2 InSAR uncertainty due to orbit error**

47 We estimated the orbital errors using the method from Fattahi & Amelung (2014) [3] using  
 48 the actual satellite orbit parameters, as shown in table S1. The precision of Sentinel-1's precise  
 49 orbits is better than 1 cm [4]. For ALOS-2, the cross-track orbit is very accurate with <8 cm in  
 50 RMS in the worst case, while the along-track orbit RMS dominates the overall 3D orbit RMS,  
 51 the latter is <25 cm [5]. Setting the azimuth correlation coefficient  $R=0.9$  (conservative) [3], we  
 52 obtained the potential contribution of the orbit error  $\sigma_r$  and  $\sigma_a$  on the InSAR secular  
 53 velocity in the range and azimuth directions, respectively, as shown in table S1 below. For  
 54 Sentinel-1, the impact of orbit error is less than 0.004 and 0.016 mm/year/100km in the range  
 55 and azimuth directions, respectively; while for ALOS-2, it is less than 0.5 and 1.1  
 56 mm/year/100km in the range and azimuth directions, respectively. We estimated the  
 57 contribution of orbital errors to the current estimated accuracy of our InSAR time series results,  
 58 as shown in Table S2. Here,  $\sigma_{total\_r}$  and  $\sigma_{total\_a}$  denote the total orbital error in the range and  
 59 azimuth directions, respectively, while  $\sigma_{total} = \sqrt{\sigma_{total\_r}^2 + \sigma_{total\_a}^2}$  represents the total potential  
 60 contribution of the orbital error.

61

62 **Table S1** Orbit error calculation parameters and results

Sensor (Mode)	$\sigma_{oh}$ [cm]	$\sigma_{ov}$ [cm]	# of acquisition /year	Total time [year]	$\theta_0$ [°]	$d\theta$ [°]	$\sigma_r$ [mm/yr /100km]	$\sigma_a$ [mm/yr /100km]
Sentinel-1 (IW)	1	1	25	9	29	7	0.004	0.016
ALOS-2 (ScanSAR)	25	8	5	7	24	7	0.30	0.66
ALOS-2 (StripMap)	25	8	10	4	4	7	0.5	1.1

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64 **Table S2** The contribution of orbital errors to the current estimated accuracy of our InSAR time series

65 results

Track	Range coverage [100km]	Azimuth coverage [100km]	$\sigma_{total\_r}$ [mm/yr]	$\sigma_{total\_a}$ [mm/yr]	$\sigma_{total}$ [mm/yr]
S1_D_USA	2.5	12	0.01	0.19	0.19
S1_D_Chile	2.5	8.5	0.01	0.14	0.14
S1_A_USA	2.5	7.5	0.01	0.12	0.12
S1_A_Chile	2.5	7.5	0.01	0.12	0.12
A2_D_USA	3.5	12	1.05	7.92	7.99
A2_D_Japan	0.5	1.5	0.25	0.28	0.38

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### 67 S3. Displacement accuracy evaluation using independent GNSS displacements

68 We use the independent GNSS displacement time series from University of Nevada at Reno  
69 [6] as reference to evaluate the accuracy of the final InSAR deformation velocity, following  
70 the steps as below:

71 **Step-1:** Select GNSS stations based on the following two criteria: (1) The GNSS station  
72 should be within the spatial extent of the InSAR dataset. (2) The GNSS time series should  
73 overlap with the InSAR time series and have more than 50 observations within this overlapped  
74 time span with the InSAR dataset.

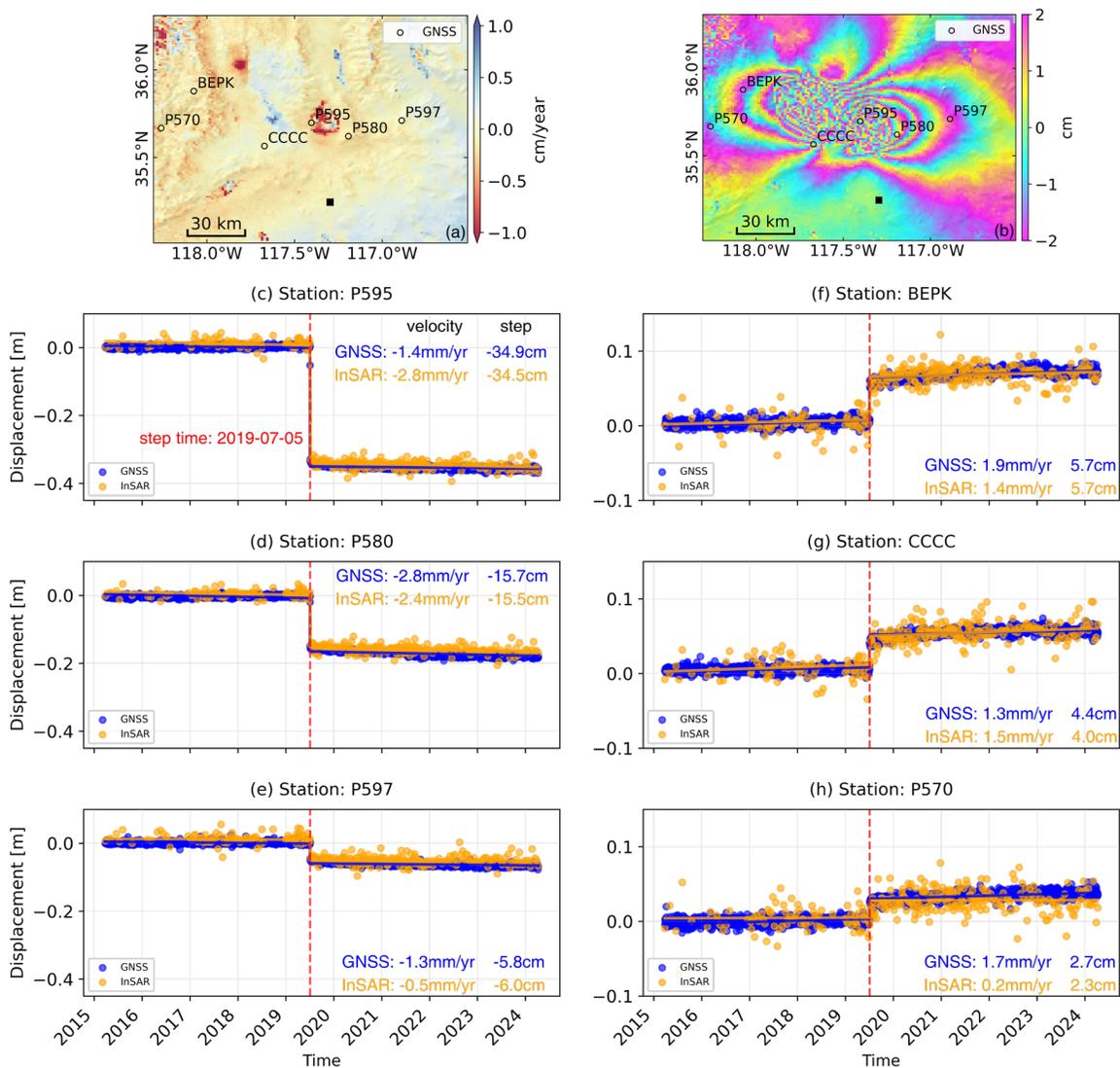
75 **Step-2:** Project the three dimensional GNSS displacement time series into the radar line-of-  
76 sight (LOS) direction. First, extract the GNSS displacement time series corresponding to the  
77 time span of the InSAR observations. Then, project the three-component (east, north, and  
78 vertical) displacements of GNSS into the LOS direction as follows:

$$79 \quad d_{los} = [d_e \quad d_n \quad d_u] \cdot \begin{bmatrix} (-1) \cdot \sin(\theta_{inc}) \cdot \sin(\theta_{az}) \\ \sin(\theta_{inc}) \cdot \cos(\theta_{az}) \\ \cos(\theta_{inc}) \end{bmatrix} \quad (S1)$$

80 where  $d_e$ ,  $d_n$ ,  $d_u$  denote the three-component (east, north, and vertical) displacements of  
81 GNSS, respectively;  $d_{los}$  represents the displacement of GNSS projected onto the LOS  
82 direction;  $\theta_{inc}$  is the incidence angle from vertical,  $\theta_{az}$  is the azimuth angle of the LOS  
83 vector from the ground to the SAR platform measured from the north with anti-clockwise  
84 direction as positive.

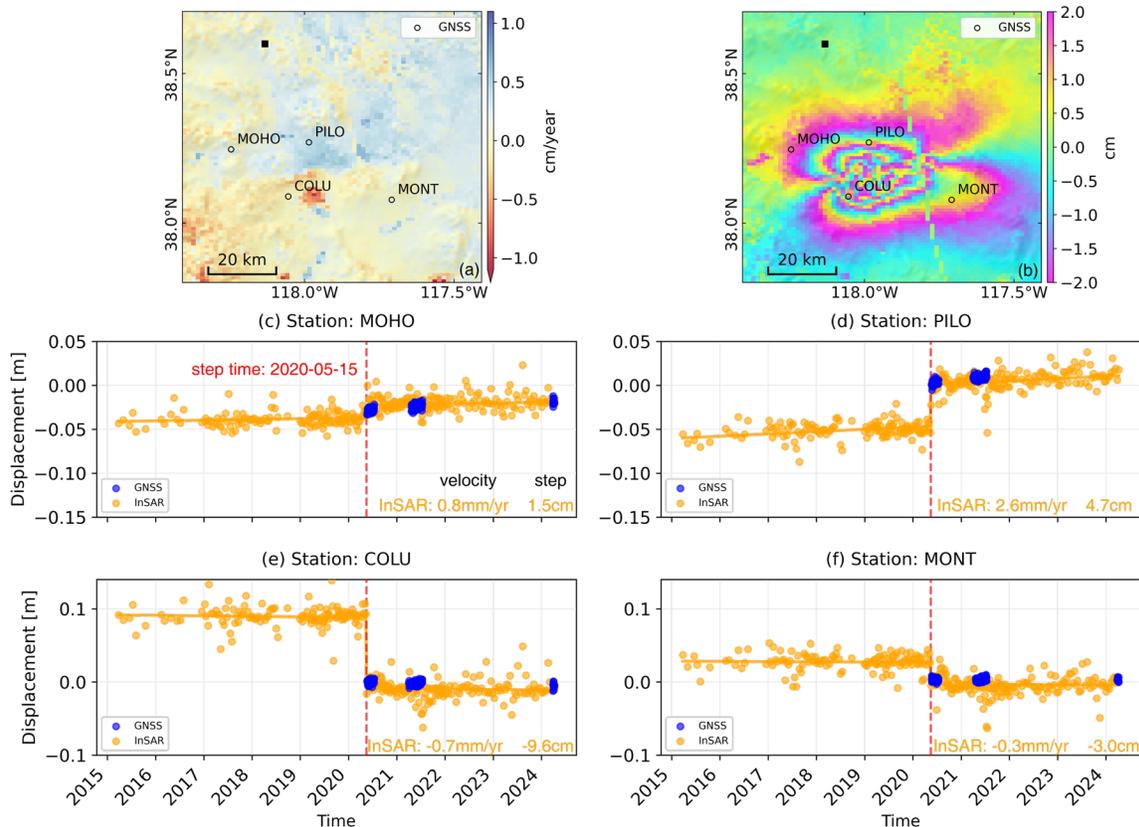
85 **Step-3:** Calculate the average velocities at each GNSS stations from GNSS and InSAR. (1)  
86 For both GNSS and InSAR, we choose a common spatial reference point at one GNSS stations  
87 for all datasets, whose ID are shown in Figs. 6-8 (except for the Sentinel-1 ascending track 149  
88 in northern Chile, where we calculate the median velocity difference between the GNSS  
89 velocities and InSAR velocities at all GNSS stations and remove this difference from all GNSS  
90 velocities). (2) We estimate a linear velocity, and a Heaviside step function for each earthquake  
91 event where applicable (Figs. S11-S13), from each GNSS station and InSAR pixels. (3) For  
92 InSAR velocities, we interpolate the regularly gridded InSAR data at each GNSS stations using  
93 a linear interpolator, to account for the potential location difference between the GNSS station  
94 and InSAR pixel center [7].

95 **Step-4:** Calculate the root mean square error (RMSE) and the coefficient of determination  
96  $R^2$  between GNSS and InSAR velocities on all stations for each InSAR dataset, following  
97 equation (7)-(8).



98

99 **Figure S11.** Comparison of InSAR and GNSS time series for Sentinel-1 ascending track 064 in western  
 100 USA account for the impact of the 2019 Ridgecrest earthquakes on 2019-07-04 and 2019-07-05 [8].  
 101 (a)-(b) Estimated linear velocity and Heaviside step function at 2019-07-05. The black square denotes  
 102 the reference point at GNSS station P592. The estimated step function is re-wrapped into  $[-2, 2)$  cm for  
 103 display. (c)-(h) InSAR and GNSS time series for all GNSS stations within  $\sim 70$  km from the earthquake  
 104 epicenter.



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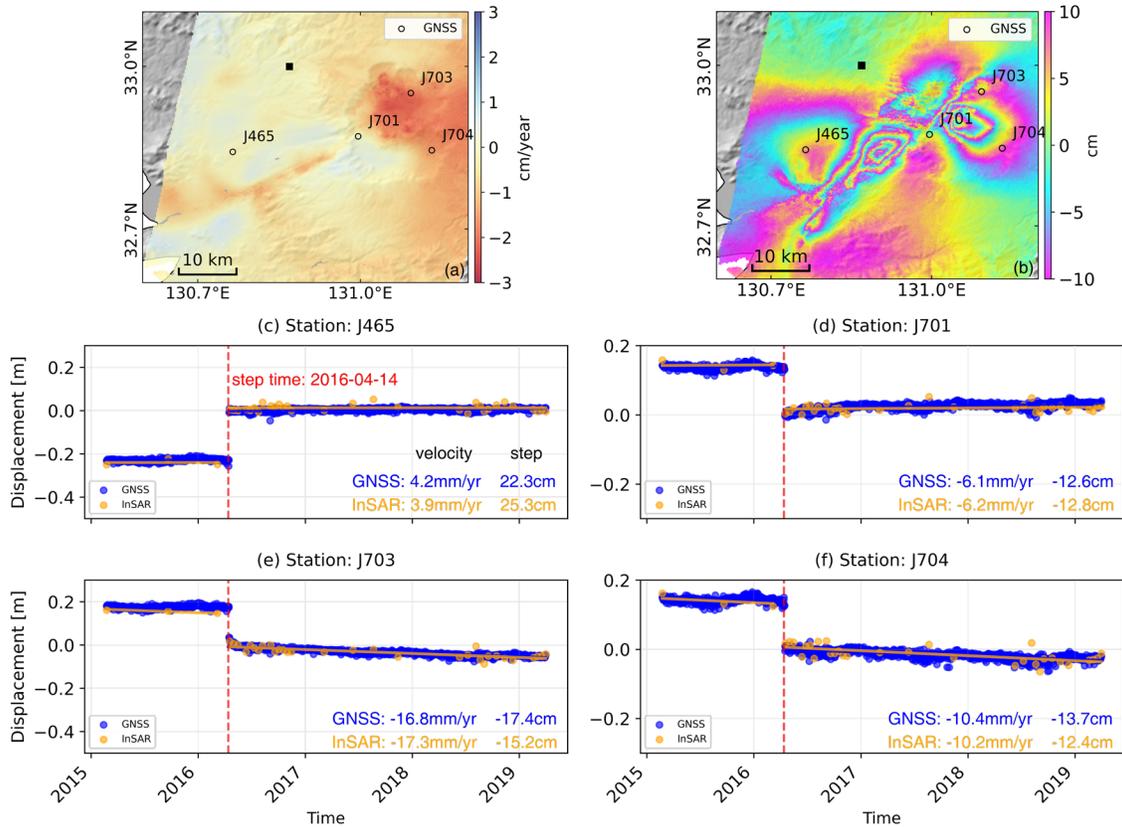
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**Figure S12.** Comparison of InSAR and GNSS time series for Sentinel-1 ascending track 064 in western USA account for the impact of the 2020 Monte Cristo Range earthquake on 2020-05-15 [9]. (a)-(b) Estimated linear velocity and Heaviside step function at 2020-05-15. The black square denotes the reference point at GNSS station CALA. The estimated step function is re-wrapped into  $[-2, 2]$  cm for display. (c)-(f) InSAR and GNSS time series for all GNSS stations within  $\sim 40$  km from the earthquake epicenter. According to the GNSS screening criteria described in section S3, the GNSS time series does not have more than 50 observations within the span with the InSAR dataset., so no GNSS comparison is performed.



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**Figure S13.** Comparison of InSAR and GNSS time series for Sentinel-1 ascending track 064 in western USA account for the impact of the 2016 Kumamoto earthquake on 2016-04-14 [10]. (a)-(b) Estimated linear velocity and Heaviside step function at 2016-04-14. The black square denotes the reference point at GNSS station G070. The estimated step function is re-wrapped into  $[-10, 10]$  cm for display. (c)-(f) InSAR and GNSS time series for all GNSS stations within  $\sim 30$  km from the earthquake epicenter.

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